

### Double Asteroid Redirection Test (DART) - NASA's First Planetary Defense Mission

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## COMESSAND ASTEROIDS **CAN STRIKE OUT OF** OWHERE BUSINESS INSIDER

https://www.businessinsider.com/no-one-saw-chelyabinsk-meteor-over-russia-coming-could-happen-again-2016-4



#### Planetary Defense: Mitigation of Asteroid Hazards, a Global Concern

Small asteroids hitting Earth, 1988–2019

#### Fireballs Reported by US Government Sensors

(1988-Apr-15 to 2019-Jul-23; limited to events  $\geq 0.1$ kt)



https://cneos.jpl.nasa.gov/fireballs/

Alan B. Chamberlin (JPL/Caltech)

#### **Chelyabinsk-sized**

Diameter: ~20 m ~500 kilotons TNT 2013, every few decades

#### **Tunguska-sized**

Diameter: 60–190 m ~5 megatons TNT 1908, every few centuries

#### Chicxulub-sized "dinosaur killer"

Diameter: 10–15 km 65 million years ago



DART



[CENTER FOR NEAR EARTH OBJECT STUDIES]

#### **SEARCH, DETECT & TRACK**

K

[GROUND-BASED & SPACE-BASED OBSERVATIONS, IAWN]

#### MITIGATE [dart, fema exercises]

### PLANETARY **DEFENSE**

IAU

Planet

Minos

#### CHARACTERIZE

[NEOWISE, GOLDSTONE, ARECIBO, IRTF]

NASA Planetary Defense Coordination Office (PDCO): established 2016 PLAN & COORDINATE

[SMPAG, PIERWG, DAMIEN IWG]

#### *Top Priority for a Mitigation Mission*

Defending Planet Earth (2010) National Academy of Sciences

<u>Recommendation:</u> "the first priority for a space mission in the mitigation area is an experimental test of a kinetic impactor"

DART is the first kinetic impact test at a realistic scale for planetary defense

#### Mitigation Techniques for Potentially Hazardous Asteroids (PHA)



DART

#### DART – Double Asteroid Redirection Test

The first mission to demonstrate asteroid deflection with a kinetic impactor



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Launch July 22, 2021



**IMPACT: September 30, 2022** 

LICIACube (Light Italian Cubesat for Imaging of Asteroids) ASI contribution



DART Spacecraft ~600 kg arrival mass 18.8 m × 2.4 m × 2.0 m 6.65 km/s closing speed

**Didymos-B** 163 meters 11.92-hour orbital period 65803 Didymos (1996 GT)

1,180-meter separation between centers of A and B

**Didymos-A** 780 meters, S-type

2.26-hour rotation period

#### **Earth-Based Observations**

0.07 AU range at impact Predicted ~10-minute change in binary orbit period

- Target the binary asteroid Didymos system
- Impact Didymos-B and change its orbital period
- Measure the period change from Earth



#### **Didymos:** The Ideal Target

It allows a deflection demonstration on an asteroid of the relevant size by changing its orbital period by ~1% about the larger asteroid.

In 2022, the result can be observed from Earth-based telescopes.







DART-1: DART shall intercept the secondary member of the binary asteroid (65803) Didymos as a kinetic impactor spacecraft during its September to October, 2022 close approach to Earth.

DART-2: The DART impact on the secondary member of the Didymos system shall cause at least a 73-second change in the binary orbital period.

DART intercepts Didymos B, designing the impact to maximize the deflection to the orbit of Didymos B about Didymos A





DART

DART-3: The DART project shall characterize the binary orbit with sufficient accuracy by obtaining ground-based observations of the Didymos system before and after spacecraft impact to measure the change in the binary orbital period to within 7.3 seconds (1- $\sigma$ confidence).

> Discovery Channel Telescope (credit: Lowell Observatory)

DART's investigation observes the eclipsing binary Didymos system before and after DART's impact to measure the deflection caused by DART **DART-4A:** The DART project shall use the velocity change imparted to the target to obtain a measure of the momentum transfer enhancement parameter referred to as "Beta" ( $\beta$ ) using the best available estimate of the mass of Didymos B.





DART's investigation determines and interprets the momentum transfer resulting from the DART impact

**DART-4B:** The DART project shall obtain data, in collaboration with groundbased observations and data from another spacecraft (if available), to constrain the location and surface characteristics of the spacecraft impact site and to allow the estimation of the dynamical changes in the **Didymos system** resulting from the DART impact and the coupling between the body rotation and the orbit.





DART

DART's investigation characterizes the impact site, dynamical effects, and ejecta to fully evaluate the results of DART's kinetic impactor mission

#### DART: Key Technologies

DART will demonstrate key technologies for future planetary defense missions.



#### Autonomous SMART Nav System

SMART Nav autonomously directs DART to impact Didymos-B, leveraging decades of missile guidance algorithm development and experience at APL.



#### DRACO Imager, modified from LORRI

DRACO (Didymos Reconnaissance and Asteroid Camera for Opnav) is a modification of the LORRI instrument used on the New Horizons mission to Pluto and the Kuiper Belt.



#### NEXT-C Ion Propulsion Engine

DART will be the first flight of NASA's Evolutionary Xenon Thruster-Commercial (NEXT-C) ion propulsion engine and be about 3x as powerful as the ion engine on the Dawn mission.



#### ROSA – Roll Out Solar Arrays

NASA's newly developed ROSA arrays will provide a compact form and light mass for launch and be deployed in space to reach 19 meters from tip to tip.



#### Didymos Reconnaissance and Asteroid Camera for OpNav (DRACO)

DAR

- DRACO is the DART payload
  - Derived from LORRI on New Horizons
  - IFOV is 4.96 µrad binned (2.48 µrad unbinned)
  - Full FOV is 0.29 degree
  - CMOS detector
  - Panchromatic
- DRACO is designed to support several tasks
  - Optical navigation of the spacecraft (OpNav)
  - Ensure impact with Didymos B (SmartNav)
  - Refine system properties (e.g., orbit, rotation rate, pole)
  - Constrain the location of the impact site
  - Characterize the surface of Didymos B
  - Determine the shape of Didymos B

### Know little about the object we are trying to hit

![](_page_17_Picture_1.jpeg)

Images centered on Didymos, moving through star fields. Taken from VLT in Chile, March/April 2019

![](_page_17_Picture_3.jpeg)

### Radar shape model

Preliminary shape model of the Didymos primary from combined radar and light curve data, diameter ~780 m.

![](_page_17_Picture_6.jpeg)

### ... and won't know in time to hit it

![](_page_18_Picture_1.jpeg)

![](_page_18_Figure_2.jpeg)

![](_page_18_Picture_4.jpeg)

![](_page_19_Figure_1.jpeg)

![](_page_19_Figure_2.jpeg)

![](_page_20_Figure_1.jpeg)

![](_page_20_Figure_2.jpeg)

![](_page_21_Figure_1.jpeg)

![](_page_21_Figure_2.jpeg)

![](_page_22_Figure_1.jpeg)

![](_page_22_Figure_2.jpeg)

Didymos B will fill ~60% of the camera frame

Itokawa at 50 cm/pix

Only "sort of know" where the object is

![](_page_23_Picture_1.jpeg)

Needle in the haystack ... and you have to thread that needle

![](_page_23_Figure_3.jpeg)

![](_page_23_Picture_4.jpeg)

![](_page_24_Picture_0.jpeg)

#### DART - Through the Eyes of DRACO Small-body Maneuvering Autonomous Real-Time Navigation SMART Nav)

DART

|   | Category                      | Parameter  |
|---|-------------------------------|--|
| oon   | Mass<br>Properties            | CM position error  |
|   |                               | Inertia principal axis misalignment                      |
|   |                               | Inertia principal moment error                           |
|   |                               |  |
| Vary the Spacecraft                                 | Mass<br>Property<br>Knowledge | CM position knowledge error                              |
| <b>y</b> 1  |                               |  |
| ulations on high fidelity<br>(hardware & emulations | IMU                           | Gyro bias per axis (post-cal)                            |
|   |                               | Gyro scale factor error per axis (post-cal)              |
|   |                               | Gyro rotational misalignment per axis (post-cal)         |
|   |                               | Gyro non-orthogonality per axis (post-cal)               |
|   |                               | Accel bias per axis (post-cal)                           |
|   |                               | Accel scale factor error per axis                        |
|   |                               | Accel rotational misalignment per axis                   |
|   |                               | Accel non-orthogonality per axis                         |
|   |                               |  |
|   | ST                            | Star tracker misalignment (post-cal)                     |
| OT SDACectatt)                                      |                               |  |
|   | LPS                           | LPS thruster coefficient scale factor error per thruster |
|   | LFS                           | Thruster misalignment                                    |

### How do we know SMART Nav works?

#### On the ground

| Row names                        |
|----------------------------------|
| Orbit Phase Angle (deg)          |
| Target Shane File                |
| Primary Shape File               |
|                                  |
| Primary Albedo                   |
| Target Badius (m)                |
| Target Radius (m)                |
| Primary Radius (m)               |
| Target Spin Rate (period, hrs)   |
| Target Orbit Radius (km)         |
| 2nd Moon Present Flag            |
| 2nd Moon Radius (m)              |
| 2nd Moon Orbit Radius (km)       |
| 2nd Moon Shape File              |
| 2nd Moon Albedo                  |
| 2nd Moon Spin Period             |
| 2nd Moon Orbit Period            |
| 2nd Moon Orbit Phasing (deg)     |
| 2nd Moon Orbit Inclination (deg) |

![](_page_25_Figure_4.jpeg)

Run sim testbeds

![](_page_25_Picture_6.jpeg)

![](_page_25_Picture_7.jpeg)

![](_page_25_Picture_8.jpeg)

![](_page_25_Picture_9.jpeg)

![](_page_25_Picture_10.jpeg)

![](_page_25_Picture_11.jpeg)

![](_page_25_Picture_12.jpeg)

ID5: Kleo

![](_page_25_Picture_13.jpeg)

ID1: kw4a

ID2: kw4b

ID6: Rashalom

ID7: sphere 1

ID4: Eros

ID3: Mithra

![](_page_25_Picture_22.jpeg)

### How do we know SMART Nav works?

![](_page_26_Picture_1.jpeg)

#### On the ground

#### Goal: Hit the center

![](_page_26_Figure_4.jpeg)

15 meter (1- $\sigma$ ) miss distance is satisfied. All cases hit within 45 meters of aimpoint

![](_page_26_Picture_6.jpeg)

### Testing SMART Nav

In Flight:

23 and 60 days after launch and again 100 days prior to impact

### Practice on moons of Jupiter

Jupiter's Opposition Planet Jupiter with Galilean's moon IO, Ganymede, Europa & Callisto 4/7/2017 - 8:30pm

> Europa Ganymede GRS

> > Meade LX90, Celestron Neximage 5, Orion Sirlus EQ-G

Callisto

Do image processing, guidance and navigation, and send data to Earth real time

![](_page_27_Picture_8.jpeg)

Fernando Roquel Torres Caguas, Puerto Rico

#### Prerequisite to Impact: Don't run out of fuel at the end

![](_page_28_Figure_1.jpeg)

![](_page_28_Picture_2.jpeg)

Most of the fuel is used early on in the mission

Early simulations show that we maneuver ~65% of time Terminal Phase (last 4 hours)

### Prerequisite to Impact: Image that asteroid!

It sure won't be happening after impact...

![](_page_29_Picture_2.jpeg)

Point this camera (telescope) at Didymos

# Solar arrays are challenging

# Light Italian CubeSat for Imaging Asteroids

![](_page_30_Picture_1.jpeg)

| LICIACube<br>Goals                                      | Description   |
|---|---|
| 1- Impact<br>ejecta plume<br>evolution                  | Obtain multiple (at least three)<br>images of DART impact ejecta plume<br>over a span of times and phase<br>angles, to allow estimation of plume<br>density structure |
| 2- Impact<br>crater                                     | Obtain multiple (> 3) images of<br>DART impact site having sufficient<br>resolution to allow measurements of<br>impact crater size and morphology                     |
| 3 - Non-impact<br>hemisphere                            | Obtain multiple (at least three)<br>images of the non-impact<br>hemisphere of Didymos B   |
| 4 - Color/<br>spectral<br>imaging<br>plume &<br>Didymos | Obtain images of the ejecta plume<br>and of asteroid target to characterize<br>color and spectral variations  |

Capable 6U CubeSat provided by Agenzia Spaziale Italiana (ASI)

Based on Argomoon CubeSat that will be flying on EM-1 mission (first flight of SLS in 2020)

Two cameras (goal of 2 m/ pixel resolution imagery)

![](_page_30_Picture_6.jpeg)

Current conops includes flyby of Didymos ~ 5 min after DART impact and downlinking data after event

![](_page_30_Picture_8.jpeg)

![](_page_30_Picture_9.jpeg)

### Ground Software Heritage

![](_page_31_Picture_1.jpeg)

![](_page_31_Picture_2.jpeg)

![](_page_31_Picture_3.jpeg)

![](_page_31_Picture_4.jpeg)

![](_page_31_Picture_5.jpeg)

#### Ground Software Mission Operations Center (MOC\*) Configurations The same set of software is used in each configuration

#### Mini-MOC / Hardware-In-The-Loop Simulator

- Supports local test of a spacecraft processor subsystem and avionics via a single computer
- Provides primary user interface for the testbed (commanding and telemetry)
- Local archive

#### I&T MOC

- Supports spacecraft I&T, allowing users to send commands to and receive telemetry from the spacecraft under test and GSE
- Central archive in the OPS DMZ
- Other users may view status and telemetry information simultaneously from client workstations connected to controlling I&T MOC workstation and in the OPS DMZ

#### Flight MOC

- Supports mission simulations, launch and operations, including via ground stations
- Central archive in the OPS DMZ
- Other users may view status and telemetry information simultaneously from client workstations connected to controlling Flight MOC workstation and in the OPS DMZ

![](_page_32_Picture_13.jpeg)

![](_page_32_Figure_14.jpeg)

![](_page_32_Figure_15.jpeg)

\*Mission Operations Center (MOC) =~ Satellite Operations Center

DART

### Typical Ground System Architecture

![](_page_33_Figure_1.jpeg)

![](_page_33_Picture_2.jpeg)

### Typical Ground Software Context

![](_page_34_Figure_1.jpeg)

### Mission / System Independent Architecture

![](_page_35_Figure_1.jpeg)

![](_page_35_Figure_2.jpeg)

- DART's Ground Software consists of 71 Computer Software Components (CSC)s
  - Most CSCs are java applications that run in separate JVM processes or application server
  - ~40 CSCs are on-line applications running with an L3Harris InControl-NG Telemetry and Commanding system
  - Remainder are off-line planning or assessment applications
- Code is shared across missions and used with mission specific configuration

High (94%) Re-use = Inexpensive DART ground solution

![](_page_35_Picture_9.jpeg)

#### Downlink Flow: Many Sources = Lots of Telemetry Virtual Channels!

![](_page_36_Figure_1.jpeg)

### Ground Software: Unique Features for DART

- Receive Telemetry Frames from DSN / ESTRACK @ 3 Mbps
- Extract image packets from frames and reconstruct image in near-real-time @ 1 image / second during Terminal Phase
- Correlate additional flight data with reconstructed image in near-real-time
- Overlay image with applicable correlated data and display to users in 2 different networks
- Separately, support timely OpNav processing of recorded images to Navigation Team

![](_page_37_Figure_6.jpeg)

#### AIDA – Asteroid Impact & Deflection Assessment

International Cooperation for an International Issue

![](_page_38_Picture_2.jpeg)

![](_page_38_Picture_3.jpeg)

![](_page_38_Picture_4.jpeg)

![](_page_38_Picture_5.jpeg)

Rendezvous: 2026

![](_page_38_Picture_7.jpeg)

### AIDA – International Collaboration for Planetary Defense

#### Hera Advances the Planetary Defense Test that DART Begins

- Hera would launch in 2024, arrive at Didymos in 2026, and conduct a 6month survey, including:
  - Directly measuring the Didymos B mass to determine momentum transfer
  - Investigating the DART impact crater depth and diameter to understand target physical property effects
- These measurements can be made four years after the DART impact without loss of value

![](_page_39_Figure_6.jpeg)

![](_page_40_Picture_0.jpeg)

![](_page_41_Picture_0.jpeg)

### JOHNS HOPKINS APPLIED PHYSICS LABORATORY