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Administration
Jet Propulsion Laboratory
California Institute of Technology

Future ground systems for scientific spacecraft

for

Ground System Architectures Workshop (GSAW2003)
Manhattan Beach, California
March 4, 2003

Gen. Eugene Tattini
Deputy Director, Jet Propulsion Laboratory
California Institute of Technology



Agenda

- **Ground systems; current situation**
- **JPL overview**
- **Future vision**



Solar system distance scales



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- **To reach Voyager 1 at 12.5 billion km, it would take:**
 - Lewis & Clark: 3.8 million years
 - JPL Director's 260Z: 14 thousand years
 - Chuck Yeager's X-1: 1500 years
 - John Glenn in the ISS: 54 years
- **For Columbus to have reached Mars by today, he should have left Spain when it was populated by Neanderthals (60,000 years ago).**
- **Alternatively, if Earth were the size of a golf ball (and humans were the size of large protein molecules), distances to other objects would be:**
 - Sun: 1600 feet (and 16 feet in diameter)
 - Mars from Sun: 2600 feet
 - Jupiter: 1.6 miles
 - Neptune: 10 miles
 - Voyager 1: 26 miles

Ground System Comparisons:

Factors Affecting Complexity



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| Some Factors Affecting GS Complexity | Example Systems | | | | |
|--------------------------------------|-----------------------|-----------------------------|-----------------------|-----------------------------|---------------------------------|
| | Commercial | DOD | NOAA | NASA Earth Orbiters | NASA Deep Space |
| Mission Customers | Mission Families | Mission Families | Mission Families | One-of-a-Kind Missions | One-of-a-Kind Missions |
| End-Data Access | Proprietary | Secure | Public | Public | Public |
| Cooperative Use of Ground Assets | Inter-Business | Inter-Service; Inter-Agency | Inter-Agency | Inter-Agency; International | Inter-Agency; International |
| Typical Link Distance | <40,000 km | <40,000 km | <40,000 km | <2,000,000 km | >>2,000,000 km |
| Typical S/C Contact Frequency | Cyclic or Continuous | Cyclic or Continuous | Cyclic or Continuous | Cyclic or Continuous | Variable |
| Two-Way Light Time Between S/C & GND | Nearly Instantaneous | Nearly Instantaneous | Nearly Instantaneous | Nearly Instantaneous | Seconds to Hours |
| Data Rates | <1 Gbps | < 1Gbps | < 1Gbps | < 1Gbps | < 10 Mbps |
| Tracking / Nav | Ranging, Doppler, GPS | Ranging, Doppler, GPS | Ranging, Doppler, GPS | Ranging, Doppler, GPS | Ranging, Doppler, VLBI, Optical |

Ground System Comparisons:

Some Cost Perspectives



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| Location | Near Earth | | | Deep Space | |
|--|------------|--------------|---------|------------|--------------|
| Mission | ALEXIS | Clementine** | SAMPLEX | CASSINI | Mars Odyssey |
| Agency | DOE | BMDO | NASA | NASA | NASA |
| Non-recurrent GS costs as a % of total mission dev. cost* | ~2% | ~2% | ~10% | ~5% | ~3% |
| Total MO&DA costs as a % of total mission cost* | ~10% | ~13%*** | ~15% | ~26% | ~25% |

*Near-Earth mission numbers derived from James R. Wertz and Wiley J. Larson, eds., *Reducing Space Mission Cost* (Torrance, California: Microcosm Press and Kluwer Academic Publishers, 1996), pp. 198-199, and 439.

**Clementine received TT&C support from NASA's Deep Space Network.

*** MO&DA percentage corresponds to planned cost – mission cut short by on-orbit software glitch.

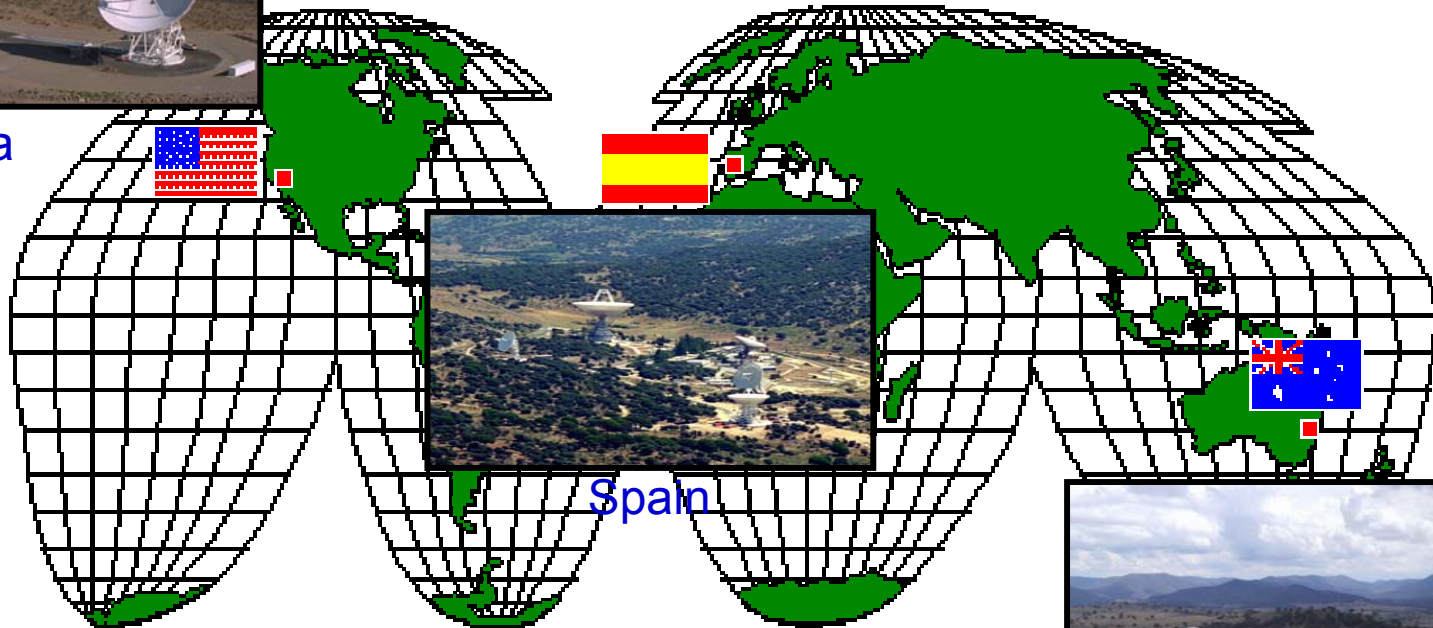
- Despite the challenges confronting these deep space missions, their non-recurrent ground system costs, as a % of their total mission development costs, are similar to that of the near-Earth missions.
- Deep space MO&DA costs as a % of total mission cost, however, tend to run higher than their near-Earth counterparts – due in large part to the increased complexity and longer duration of the missions.



Deep Space Network (DSN)



California



Spain



Australia

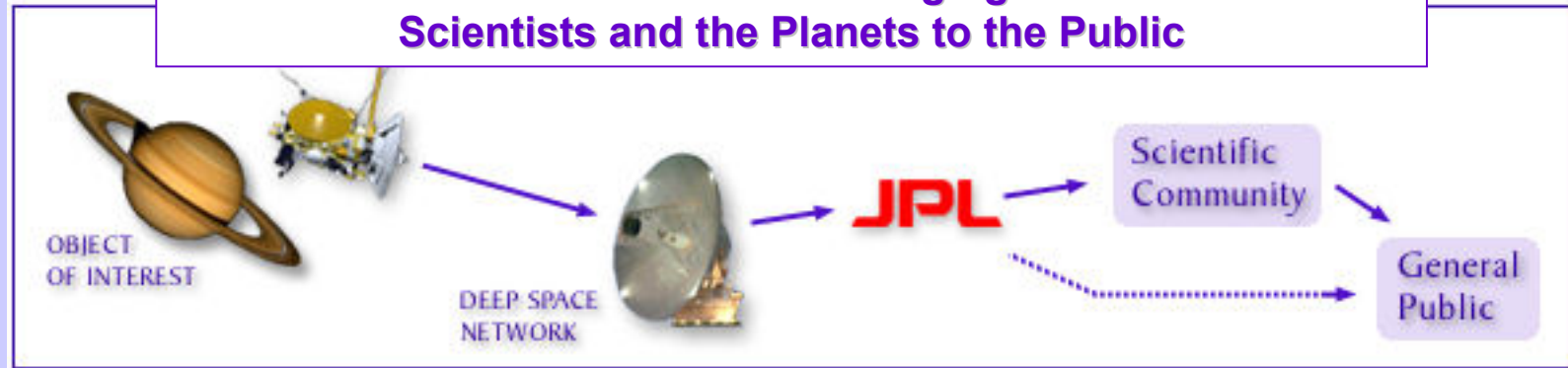
**Today's Deep Space
Communications Complexes (DSCCs)**

Serving Scientists and the Public



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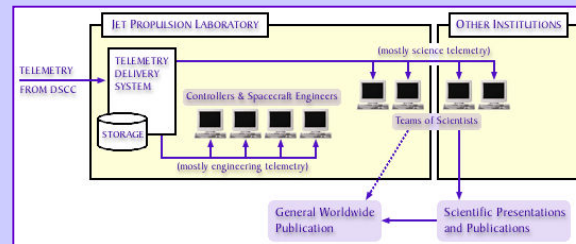
What the Network Enables: Bringing Sensors to the Scientists and the Planets to the Public



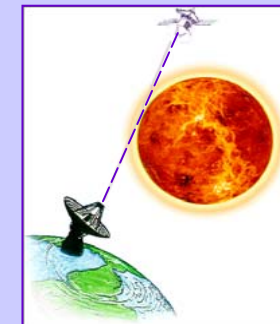
How the Network Enables It: Multi-mission Services & Tools



Telemetry & Command

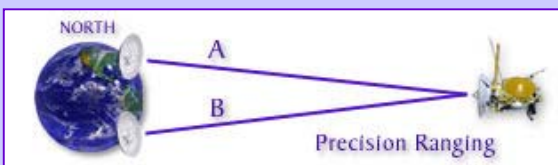


Ground Communications & Mission Data Management

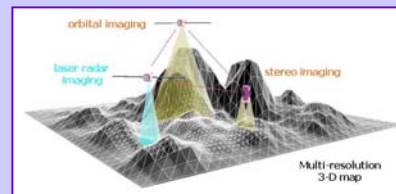


Science Services

- Radar
- Radio Astronomy
- Radio Science
- VLBI



Tracking & Navigation



Experiment Data Product Generation & Science Visualization

Ground Systems Overview



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JPL Deep Space Definition:

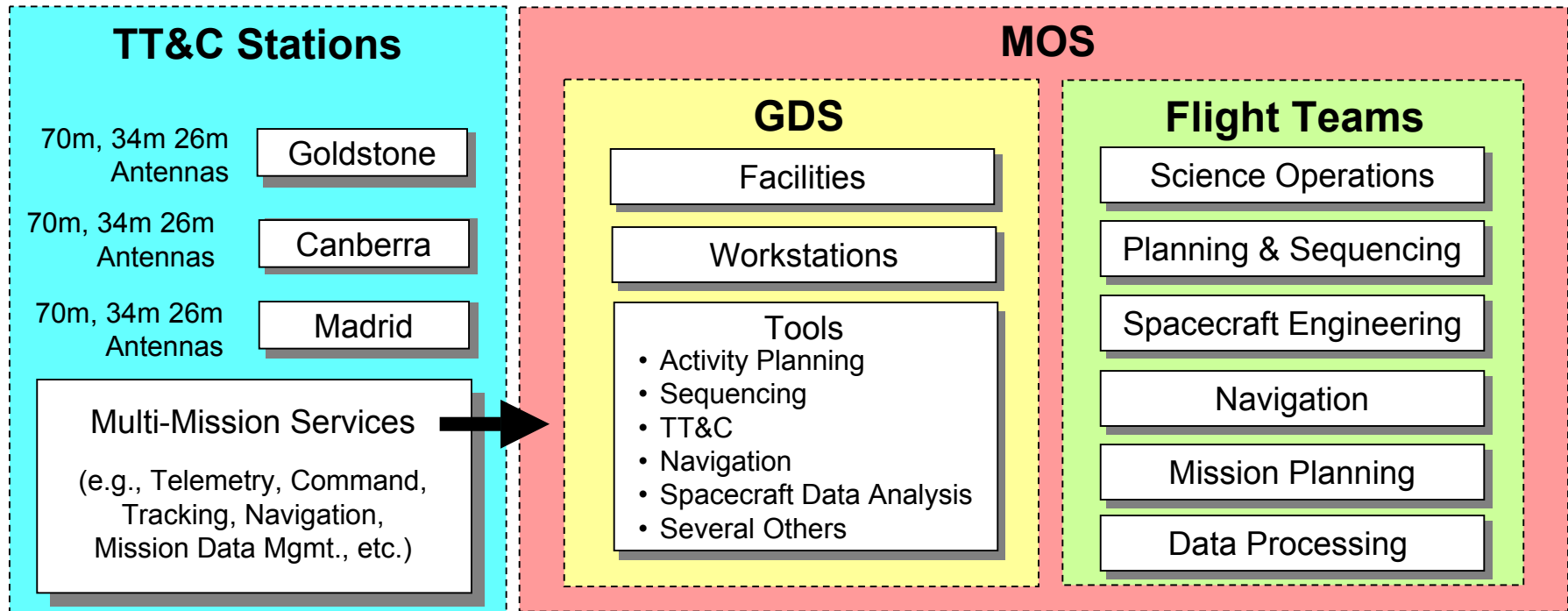
Ground System = TT&C Stations + MOS

MOS = GDS + Flight Teams

TT&C = Tracking, Telemetry, & Command

GDS = Ground Data System

MOS = Mission Operations System



Some key differences
from other ground
systems:

- Distributed operations (e.g., spacecraft ops, science ops, data acquisition, etc.)
- Each mission is unique, requiring unique tool adaptations
- Interoperability with international mission GDS's and tracking assets
- Signal-to-Noise-Ratio-constrained TT&C; long two-way light times
- Exotic tracking & navigation techniques; no GPS
- Integrated suite of multi-mission tools and services



NASA Vision and Mission

- ***NASA Vision:***
 - To improve life here;
 - To extend life to there;
 - To find life beyond.
- ***NASA Mission:***
 - To understand and protect our home planet;
 - To explore the universe and search for life;
 - To inspire the next generation of explorers as only NASA can.





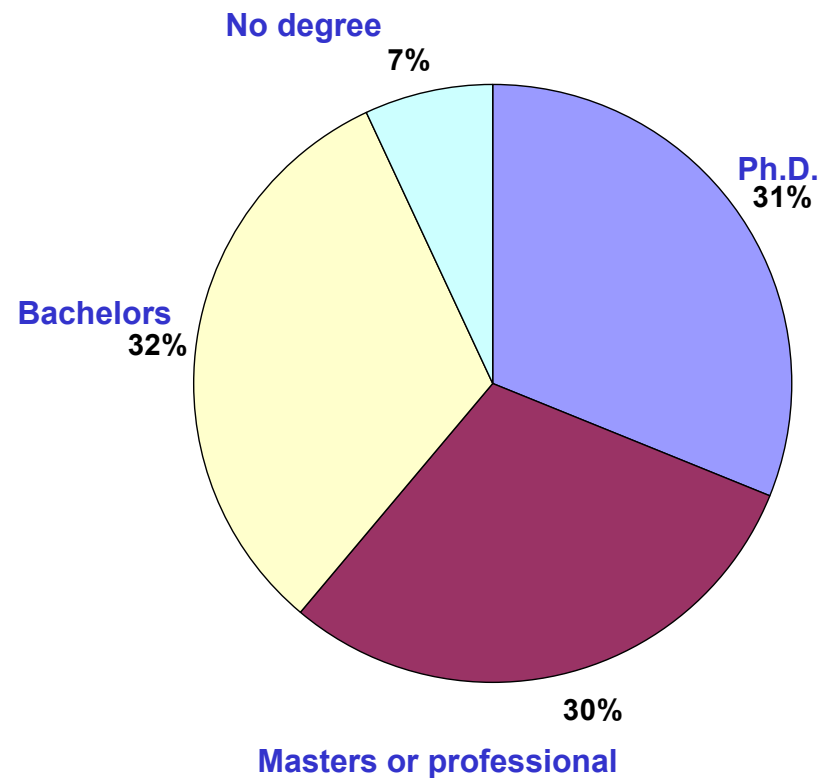
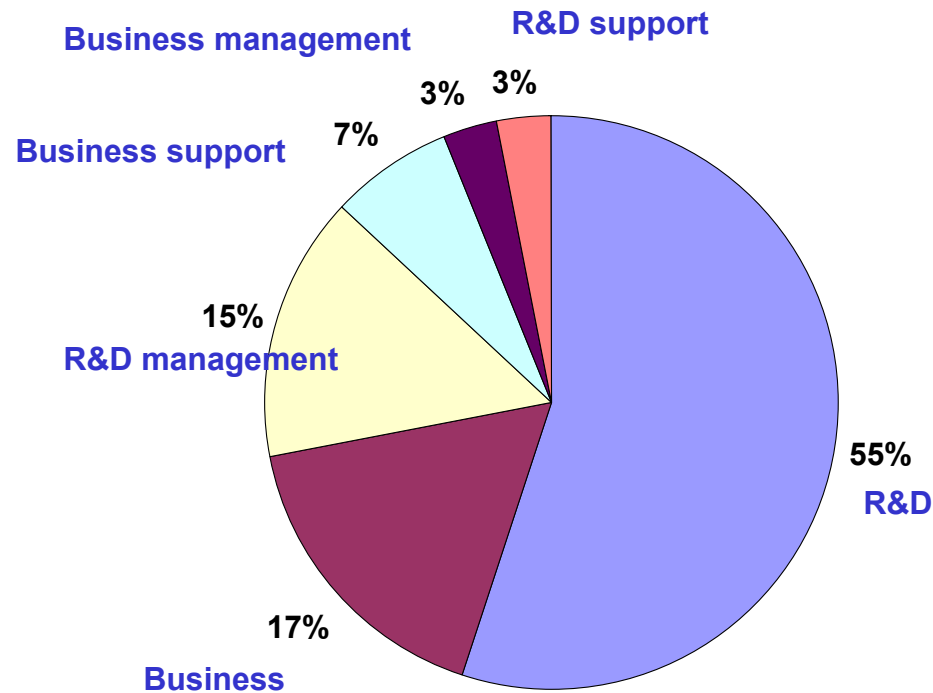
The Jet Propulsion Laboratory:

- **Has a dual character:**
 - **A unit of Caltech, staffed with Caltech employees;**
 - **A Federally-Funded Research and Development Center (FFRDC) under NASA sponsorship;**
- **Is a major national research and development (R&D) capability supporting:**
 - **NASA programs;**
 - **Defense programs;**
 - **Civil programs of national importance compatible with JPL capabilities.**



JPL staff composition by job classification and academic degree in FY2002

- Staff composition by job classification for 5175 employees
- R&D staff distribution by academic degree for 2867 employees



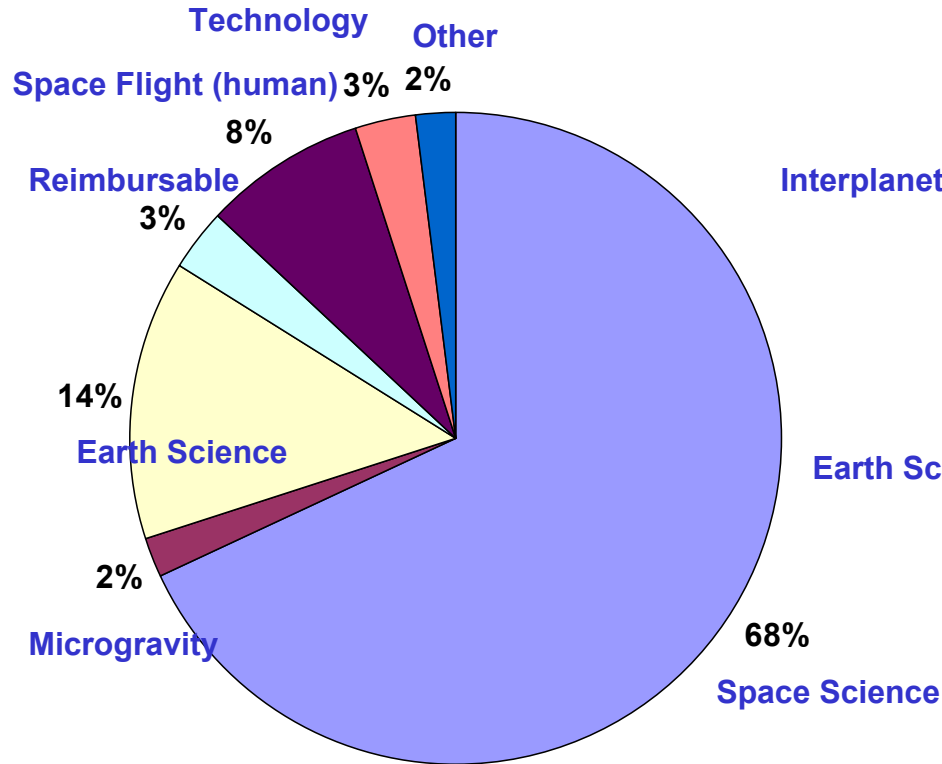
JPL funding distributions for FY02

\$1.391 billion business base

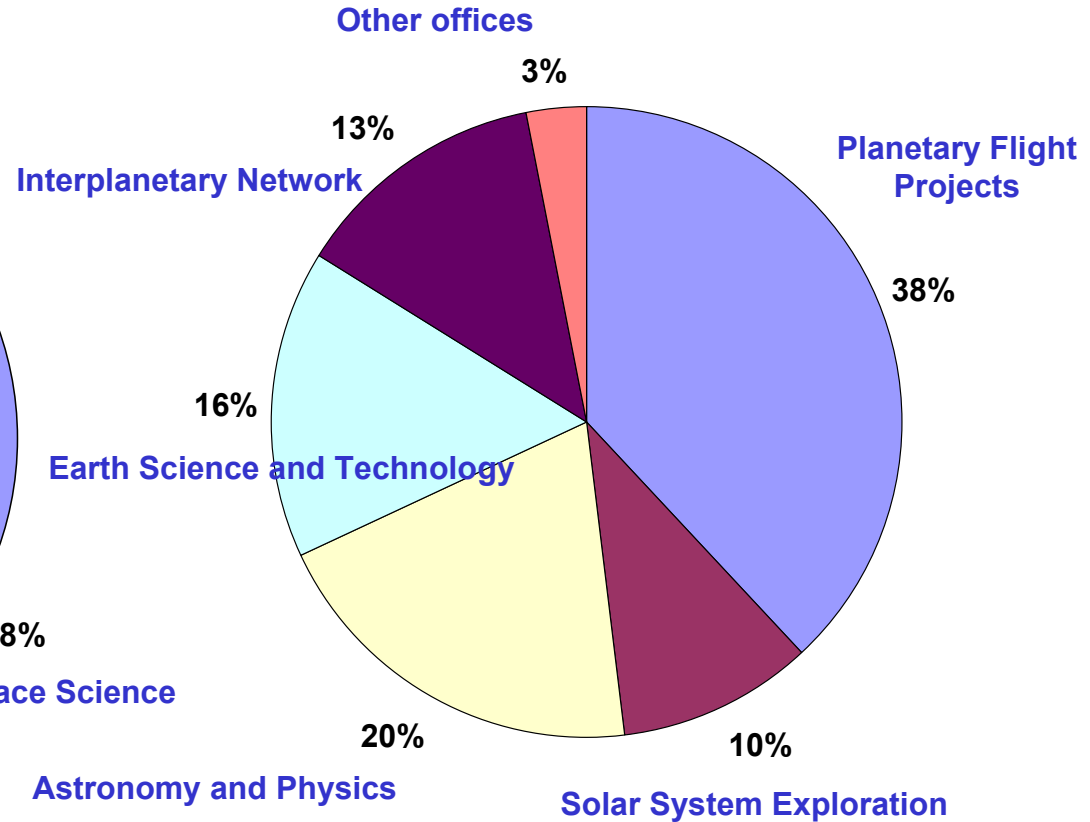


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- By NASA office or other sponsor



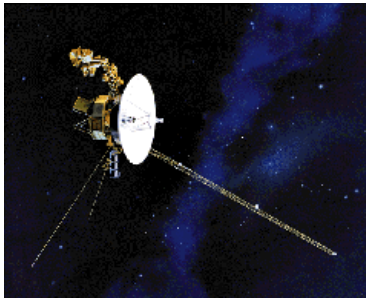
- By implementing JPL directorate



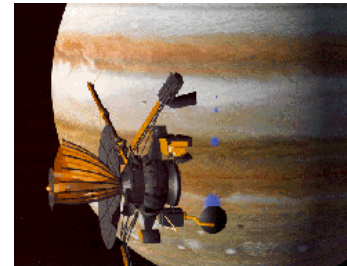
Fourteen JPL spacecraft, and three major instruments, now operating across the solar system



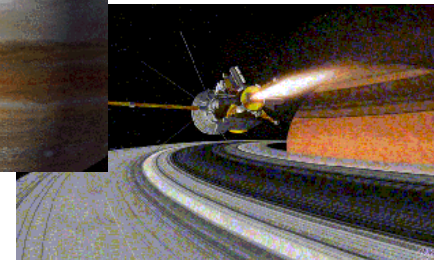
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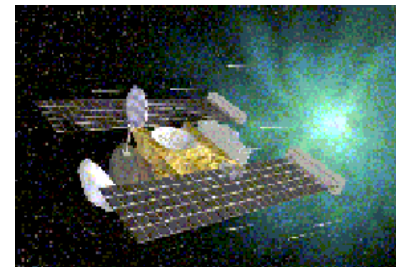
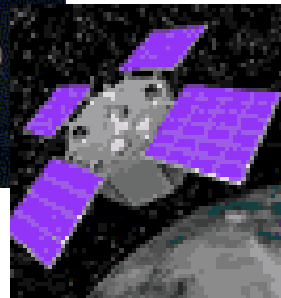
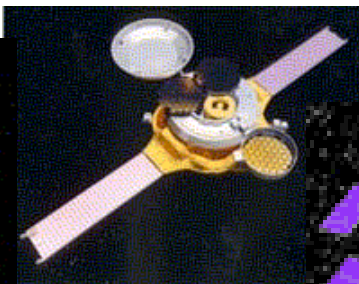
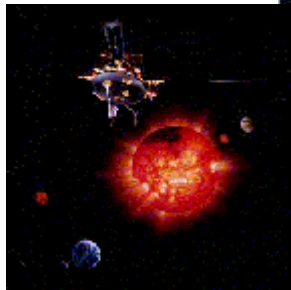
Two Voyagers on an
interstellar mission



Ulysses, Genesis, and
ACRIMSAT studying the sun



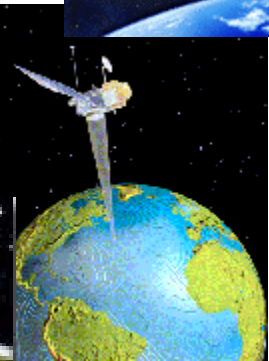
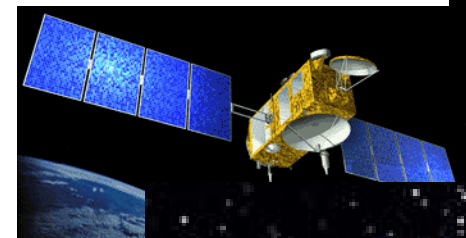
Galileo and Cassini studying
Jupiter and Saturn



Stardust returning comet dust



Mars Global Surveyor and Mars
Odyssey in orbit around Mars



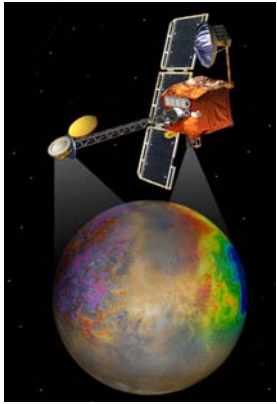
Topex/Poseidon,
Quickcat, Jason 1, and
GRACE (plus Seawinds,
MISR, and AIRS
instruments) monitoring
Earth



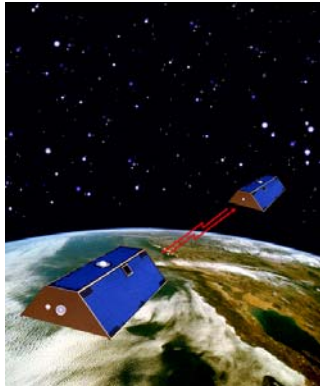
Significant recent and future events



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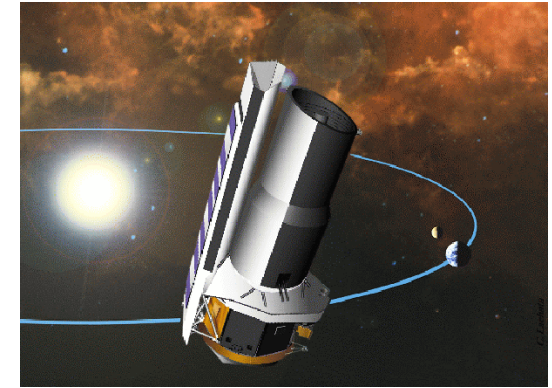
2001 Mars Odyssey
began mapping
February 2002



GRACE Earth gravity
measuring mission
launched March 17, 2002



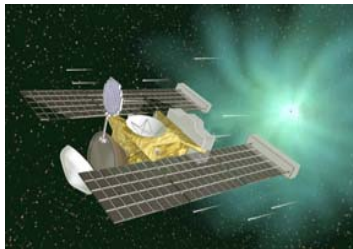
GALEX ultraviolet
observatory launch in
March 2003



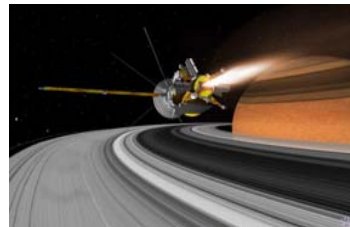
NASA infrared great
observatory SIRTF launch
in April 2003



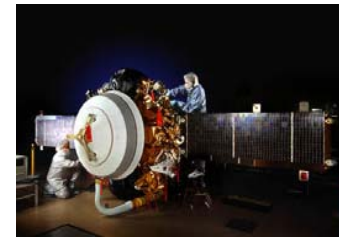
Mars Exploration Rovers
launch summer 2003,
arrive January 2004



Stardust captures
material from Comet
Wild 2 in January
2004



Cassini/Huygens
arrives at Saturn
July 2004



Genesis solar wind
sample return
September 2004



Cloudsat launch
November 12, 2004



**Hardware (and software)
designs and
implementation are
verified during the
assembly, test, and
launch operations phase.
(Mars Exploration Rover 2003
in vibration test)**



2003 - 2004: The Busiest Period in JPL's History

| | |
|-------------------|---|
| March 2003 | Galaxy Evolution Explorer (GALEX) launch |
| April 2003 | Space Infrared Telescope Facility (SIRTF) launch |
| May 30, 2003 | Mars Exploration Rover – 1 (MER-1) launch |
| June 25, 2003 | Mars Exploration Rover – 2 (MER-2) launch |
| January 2, 2004 | Stardust Encounter with Comet Wild-2 |
| January 4, 2004 | Mars Exploration Rover – 1 (MER-1) landing |
| January 25, 2004 | Mars Exploration Rover – 2 (MER-2) landing |
| January 2004 | Microwave Limb Sounder (MLS) and Tropospheric Emission Spectrometer (TES) launch on EOS-AURA |
| July 1, 2004 | Cassini Saturn orbit insertion |
| September 8, 2004 | Genesis solar wind sample return (first samples from beyond lunar orbit) |
| October 26, 2004 | First Cassini images of Titan surface |
| November 12, 2004 | Cloudsat launch |
| January 2005 | Deep Impact Launch |
| January 14, 2005 | Huygens probe Titan atmospheric entry |

In addition to the above key events:

- 7 to 12 missions in development
- 14 missions in operations

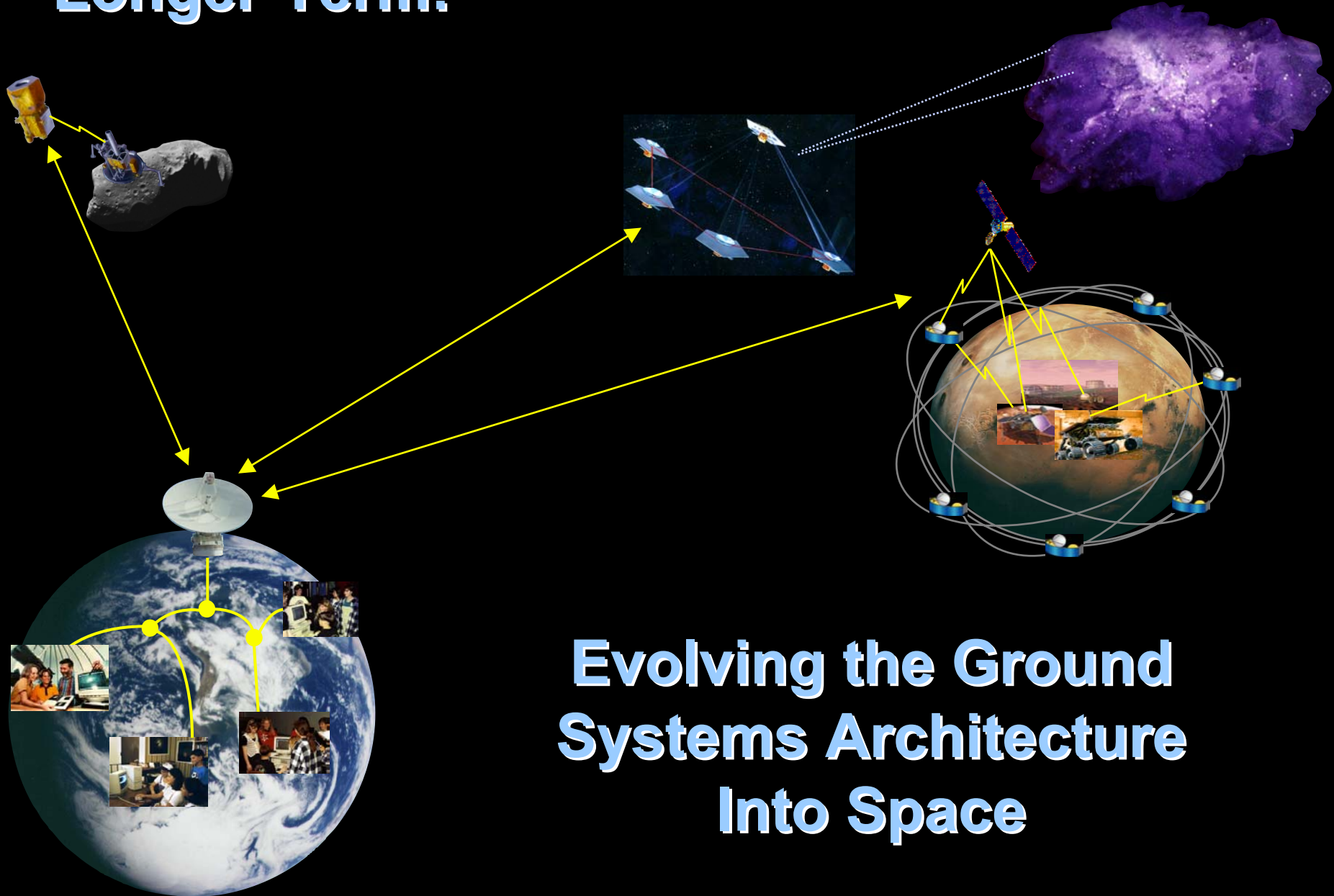


Hubble Deep Field size comparisons

- **Image:**
 - Image is of a sky region the size of Roosevelt's eye on a dime held at arm's length.
- **Galaxies in image:**
 - The smallest galaxies in the Hubble Deep Field image have a diameter seen from Earth of $1/200^{\text{th}}$ the width of a hair held at arm's length.
 - Galaxies (the dots in the Hubble Deep Field image) have:
 - ~400 billion stars
 - Mass of ~1 trillion solar masses
 - Diameters of ~1 billion billion miles
 - Distances from Earth of ~1 hundred billion trillion miles



Longer-Term:



The Changing Mission Paradigm



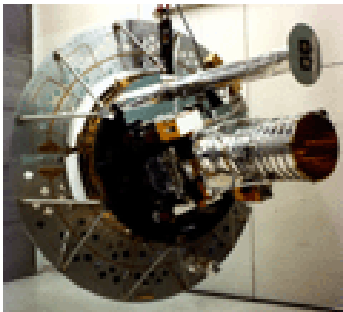
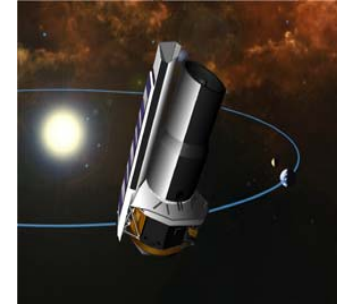
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Low-Earth-orbit
solar and
astrophysical
observatories.



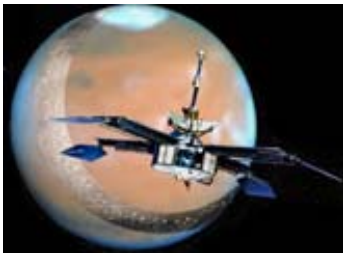
Observatories
located further
from Earth.



Single, large
spacecraft for solar
and astrophysical
observations.



Constellations of
small, low-cost
spacecraft.



Preliminary
solar system
reconnaissance
via brief flybys.



Detailed
Orbital Remote
Sensing.



In situ
exploration via
short-lived
probes.



In situ
exploration via
long-lived mobile
elements.

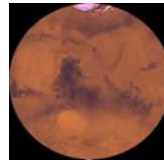




Operational Challenges

Fundamental Obstacles

- **Extreme distance** – communicating at Neptune (30 AU) is ~10 billion times more difficult than at a commercial GEO satellite distance.
- **Long Round Trip Light Times** – over 8 hours at Neptune; no “joy-sticking” possible.
- **Unique Navigation Scenarios** – small body ops, gravity assist trajectories, aerocapture/aerobraking, low-thrust propulsion, Lagrange point missions, formation flying.
- **High Launch/Delivery Cost per Unit Payload Mass** – drives need for low mass, low power flight systems.



Programmatic “Bottlenecks”

- **Deep Space Network Congestion** – compromises science return and adds risk to all missions (e.g., Mars '03-'04).
- **Limited Connectivity at Mars** – Mars science orbiters provide only limited relay communications for surface vehicles; little or no communications during many critical events.
- **Aging Assets & Insufficient Bandwidth** – ~30-year old 70m antennas; very low data rates from planets; can only map ~1% of Mars at high resolution due to data rate constraints.
- **Increasing Operations Complexity** – scientists spend more time on operations than science; more multi-element missions will increase this complexity.

Meeting the Challenges



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Build the
Deep Space
Telecom
Backbone

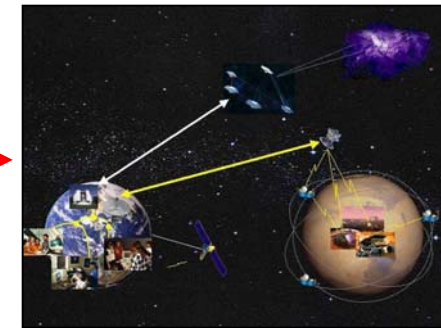
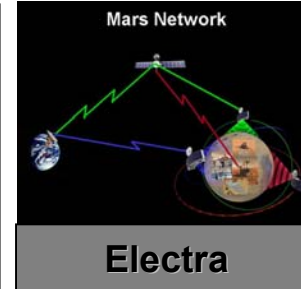
Modernize DSN
& Advance RF



Pioneer
Optical Comm



Network Space
Comm Assets

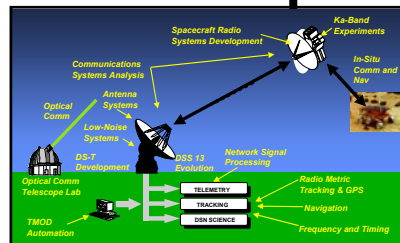


An
Interplanetary
Network

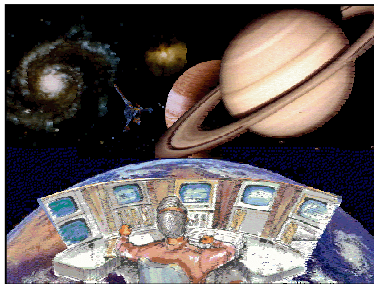
Sensors to the
Scientists

Planets to the
Public

Develop
the Tools &
Techniques
Needed to
Operate with
this
Backbone



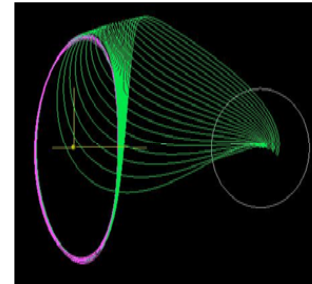
- Standards & Protocols
- S/C Comm Components
- RF & Optical Technology
- Middleware & Applications
- Mission Ops Technology



Provide Multi-Mission
Ops Systems &
Software



Revolutionize
Mission
Operations



Advance Mission
Design & Nav

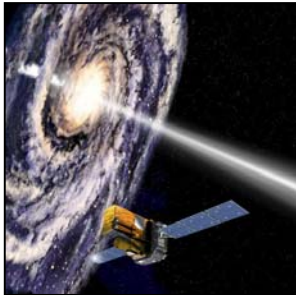


First Steps



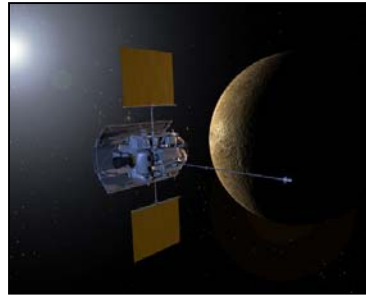
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Space Link Extension



2002: INTEGRAL

Turbo Code



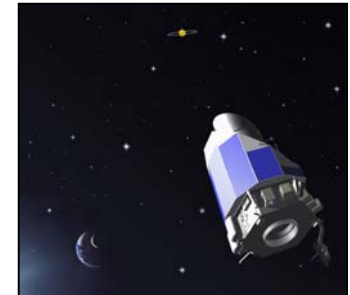
2004: MESSENGER

Ka-band (Ops Validation)



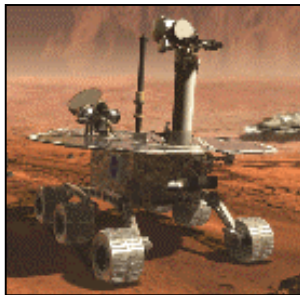
2005: MRO

Ka-band (Operational)



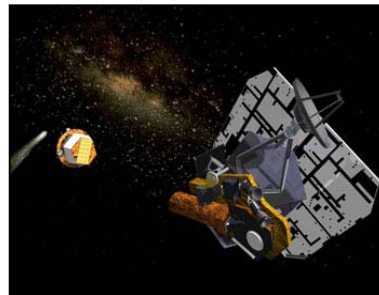
2006: Kepler

Proximity Links



2003: MER

CCSDS File Delivery Protocol



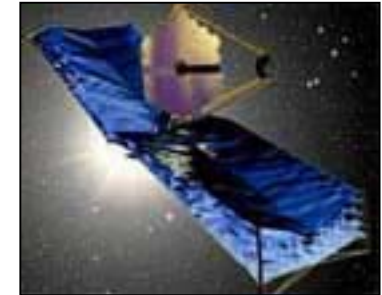
2004: Deep Impact

Higher Data Rates (>2 Mbps)



2005: MRO

Higher Data Rates (>40 Mbps)



2010: JWST
(Tentative)

Longer-Term: Large Arrays of Small Antennas



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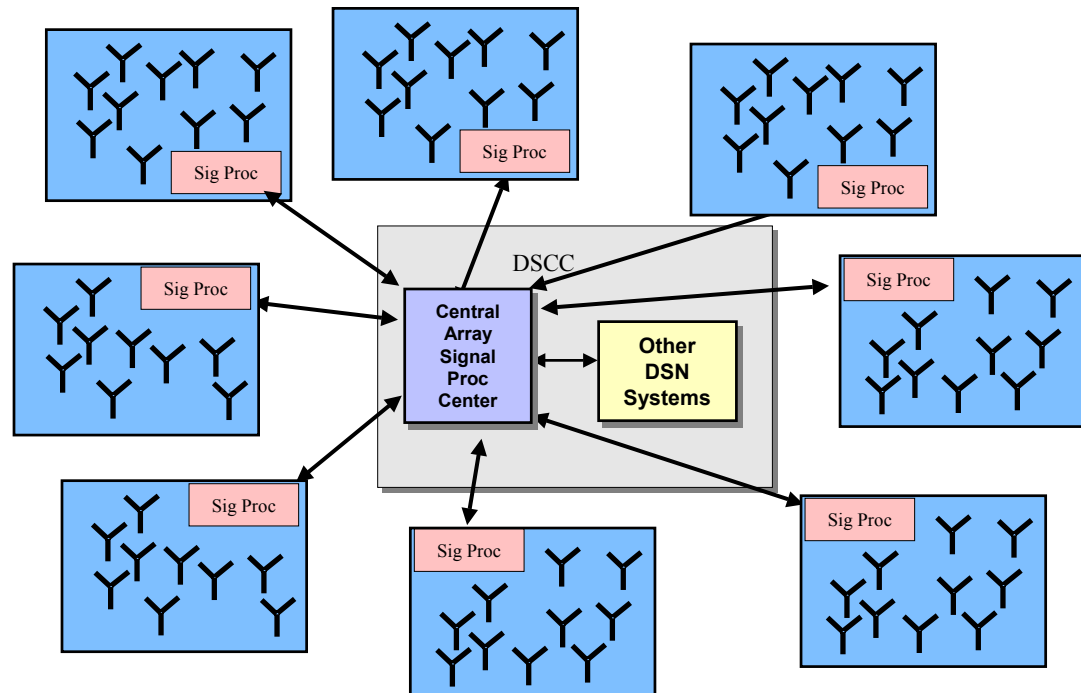
- Locations

- Large number (~3600) of small (~12m) antennas, approximately equally distributed at approximately eight sites on each of three continents.



- Processing

- Analog communication from multiple antennas in a cluster to a Local Signal Processing Center (LSPC).
 - Communication from LSPC to an Array Central Signal Processing Center (ACSPC) at a DSCC.
 - Conventional signal processing at the DSCC, with existing DSN ground communications to JPL/Customer.



Longer-Term: Optical Networks



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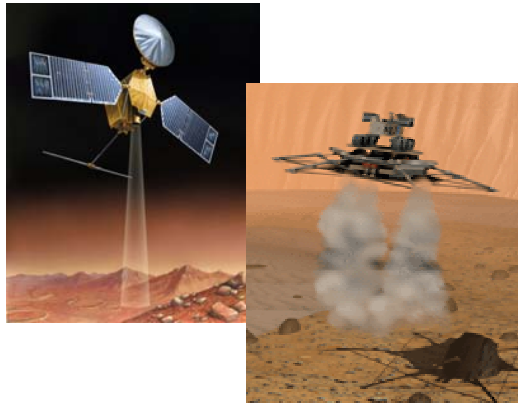
- Potential Locations
 - Linear Dispersed Optical Subnet (LDOS) - has seven stations equally spaced around the world.
 - Clustered Optical Subnet (COS) is a 9-station network with clusters of three stations every 120 deg longitude, say near the DSCCs.
 - Elevation may range from 1,000 to 3,000 meters to be free from dust.
 - Factors include geo-political realities. Oceans, unstable countries.
- Processing
 - Autonomous, with occasional on-site maintenance functions.
 - No foreseeable spectral bandwidth issues.
 - Since data rate is dependent on the capacity of the link, it will change with technology.
 - Recent analyses have predicted data rates in the 30-300 Mbps from Mars, depending on distance and the technology growth.
 - If technology developed more aggressively, these numbers could increase even further.





Our vision: JPL's legacy by 2020

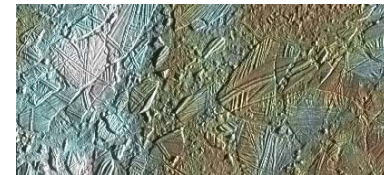
Established a continuous presence around and on the surface of Mars



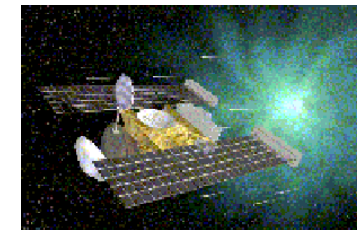
Explored Saturnian system, especially Titan, the only satellite with an organic atmosphere.



Explored Jovian satellites in detail and probed their interiors for possible life-favorable environments.



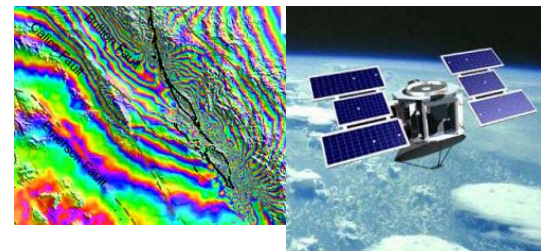
Returned first samples from other solar system bodies beyond the moon.



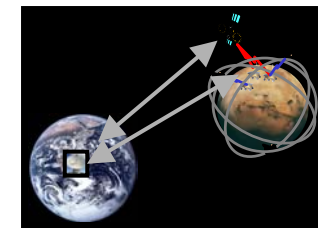
Explored the boundaries of physics to understand the forces that powered the Big Bang



Established operational capability to monitor dynamics of solid Earth and its oceans and atmosphere.



Established the Interplanetary Network, which is being commonly used by students.



Enabled efficient access to all the bodies of the solar system

